

What's Going On?

Checking In

Minds on

5 Years of Sol

Action!

The Layers of the Sun

Consolidation

Fill it In, Fill it In

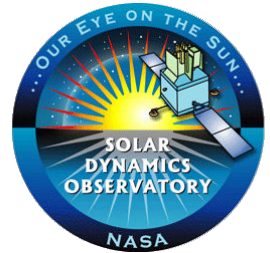
Learning Goal - I will be able to identify the key features of the Sun, and explain how the Sun affects the Earth.

Checking In

Housekeeping

Minds on

Five Years of Sol

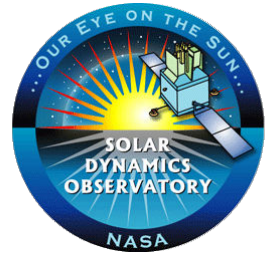


Nasa's Solar Dynamics Observatory (SDO) has been capturing images of the Sun more than once every second for the past 5 years.

On February 11, 2015 NASA released a video showcasing highlights from the last 5 years of sun watching.

Minds on

Five Years of Sol



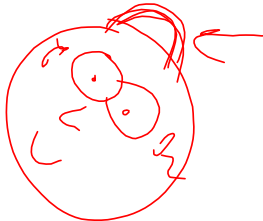
Can you name any of the things you saw in the video?

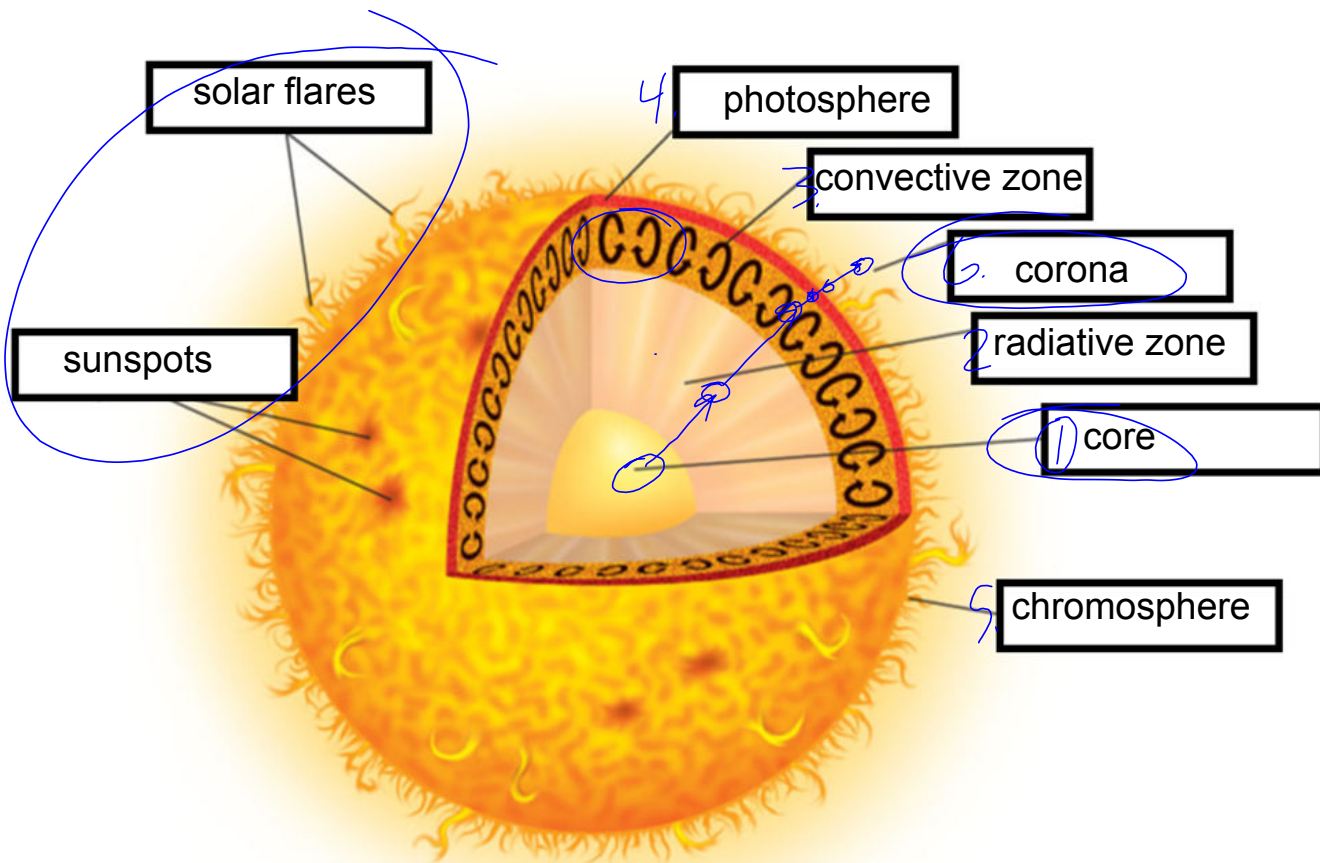
Solar Flares

Sun ☺

Mercury

Sunspots

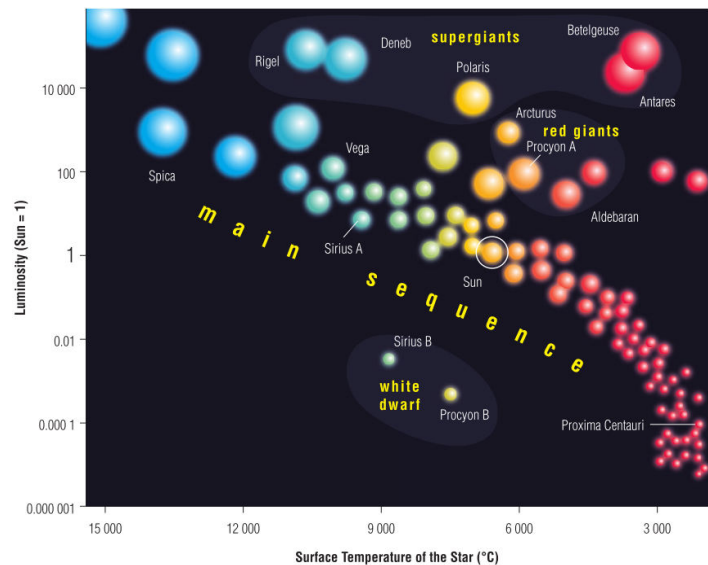




Minds on

Our Sun

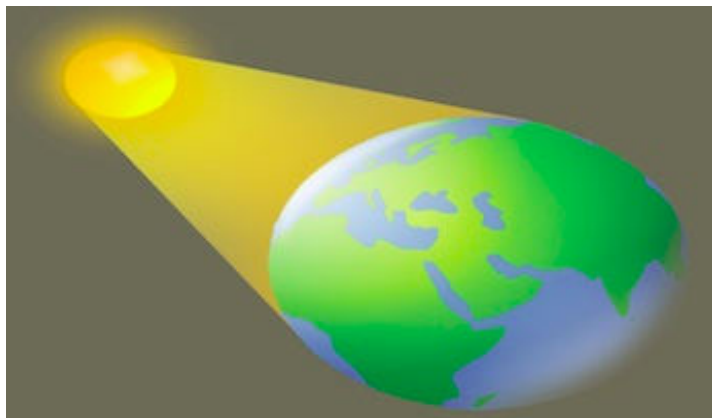
Our sun is a medium-sized star, and is the centre of the solar system. It formed inside a nebula and began shining approximately 5 billion years ago. Our Sun may continue to shine for another 5 billion years.



Minds on

Our Sun

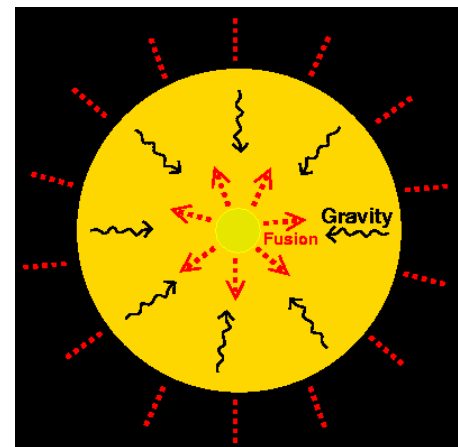
The sun emits energy in the form of heat, light
and other types of radiation including
ultraviolet radiation.



Minds on

Our Sun

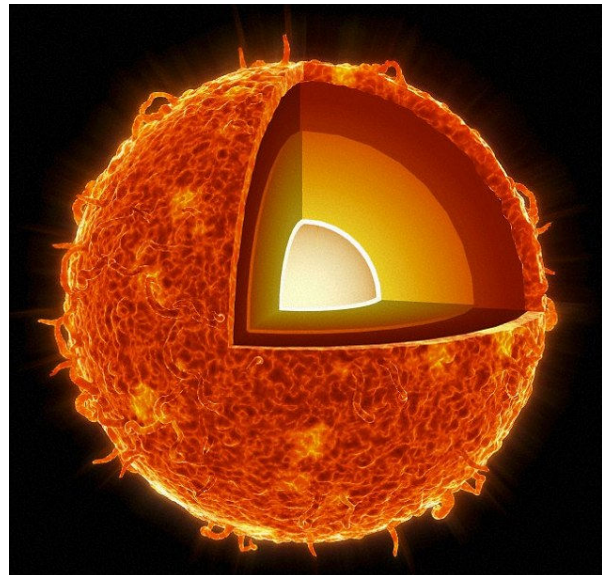
The Sun is composed mainly of hydrogen and helium, and produces energy through the process of nuclear fusion at its core.



Minds on

Our Sun

The Sun has 6 layers. Starting from the centre and moving outwards, the layers are: the core, the radiative zone, the convective zone, the photosphere, the chromosphere and the corona.



Action!

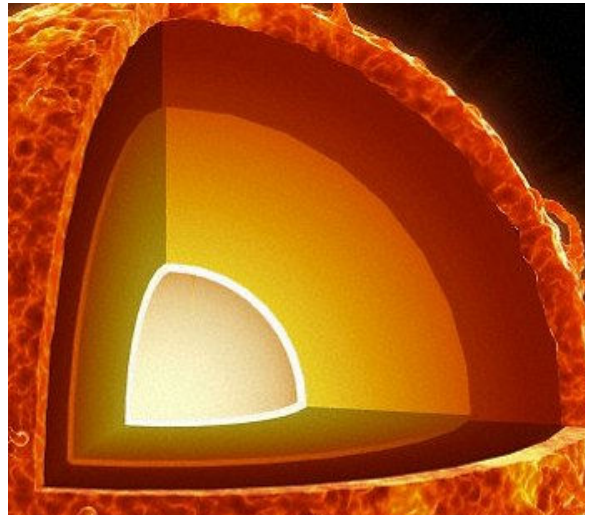
The Layers of the Sun



Action!

The Core

The innermost layer of the Sun and the site of nuclear fusion. The core is an area of very high pressure where temperatures can reach 15 million °C.



Action!

The Radiative Zone

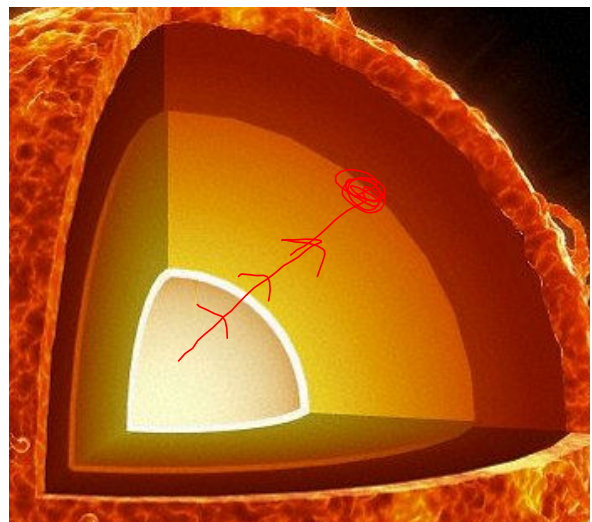
Surrounds the core and extends three quarters of the way to the surface.

Heat and light emitted from the core

move through this layer. Light takes at least

100,000 years to travel from the core

through the radiative zone.



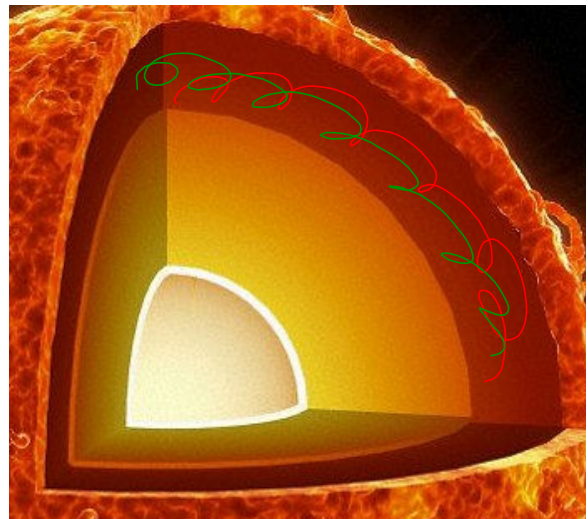
Action!

The Convective Zone

Outside the radiative zone. plasma

circulates here and bubbles towards the surface, carrying

energy.



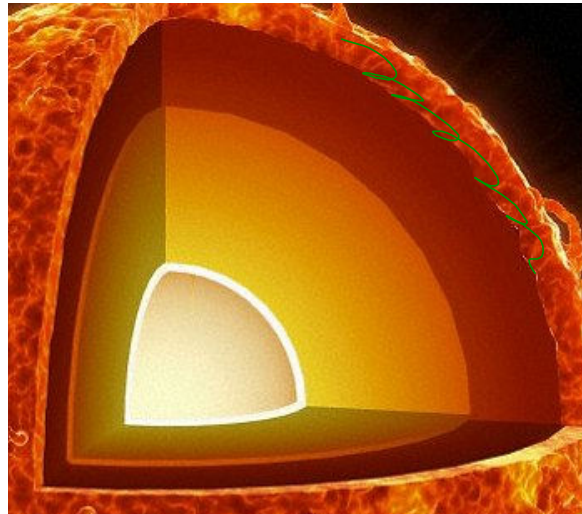
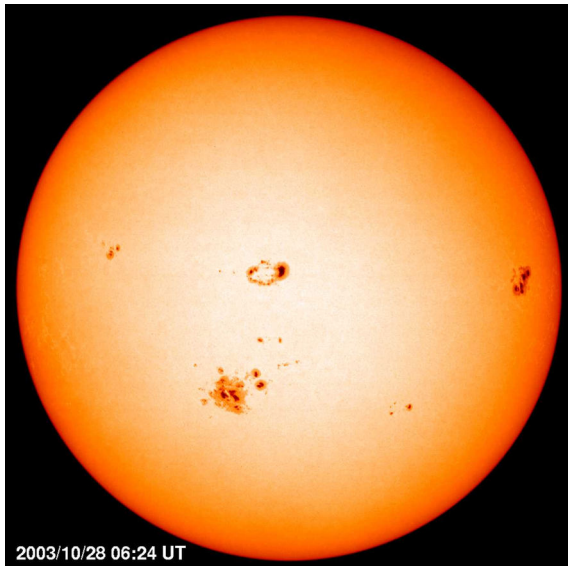
Action!

The Photosphere

Considered the boundary between the inside and outside of the Sun.

This is the part that we see from Earth and gives the sun its yellow colour.

The lowest temperatures at this layer are 2500 °C.



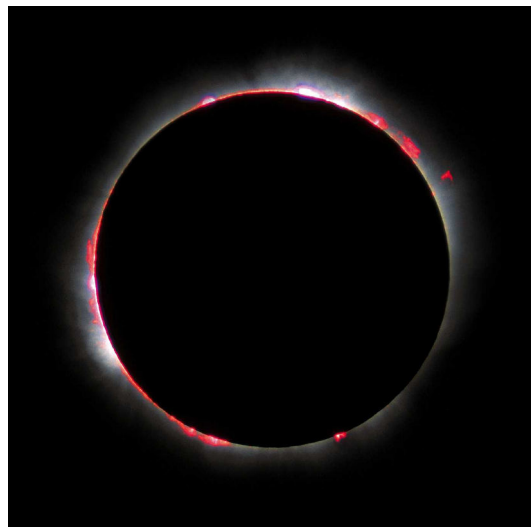
Action!

The Chromosphere

The thin layer above the photosphere.

Red in colour but only seen during a

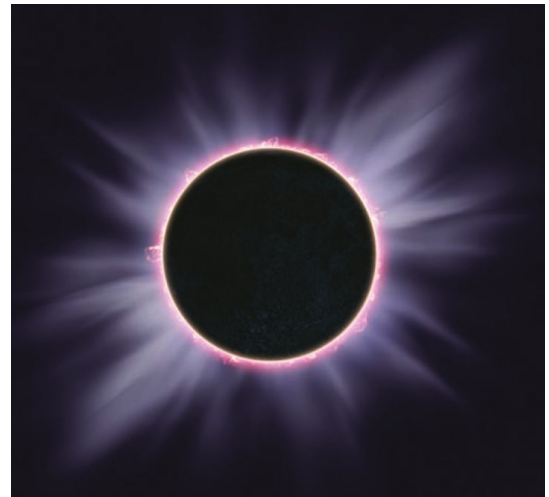
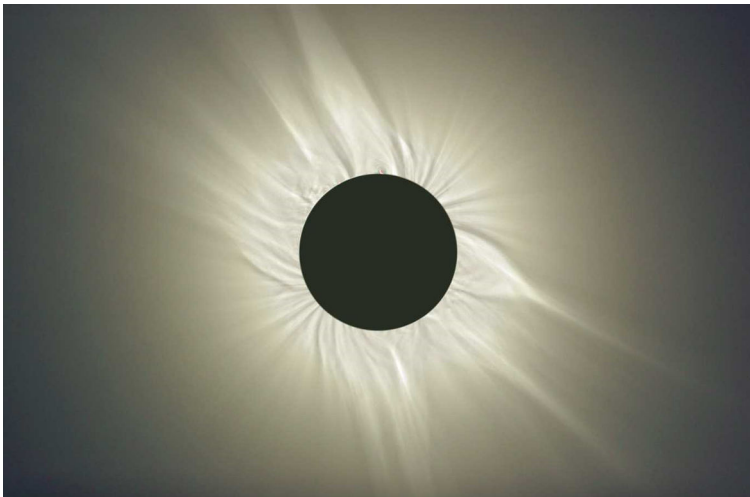
solar eclipse.



Action!

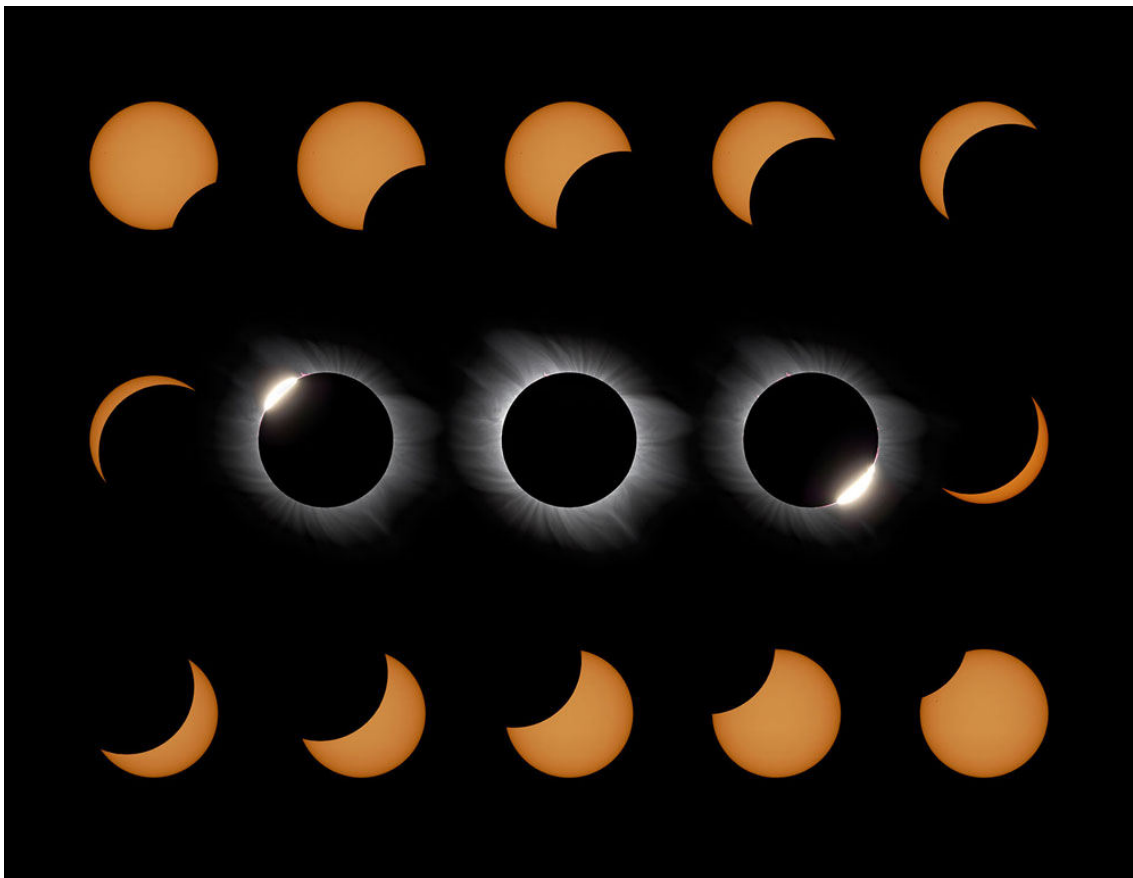
The Corona

The outermost layer, extending
millions of kilometers beyond the
chromosphere.



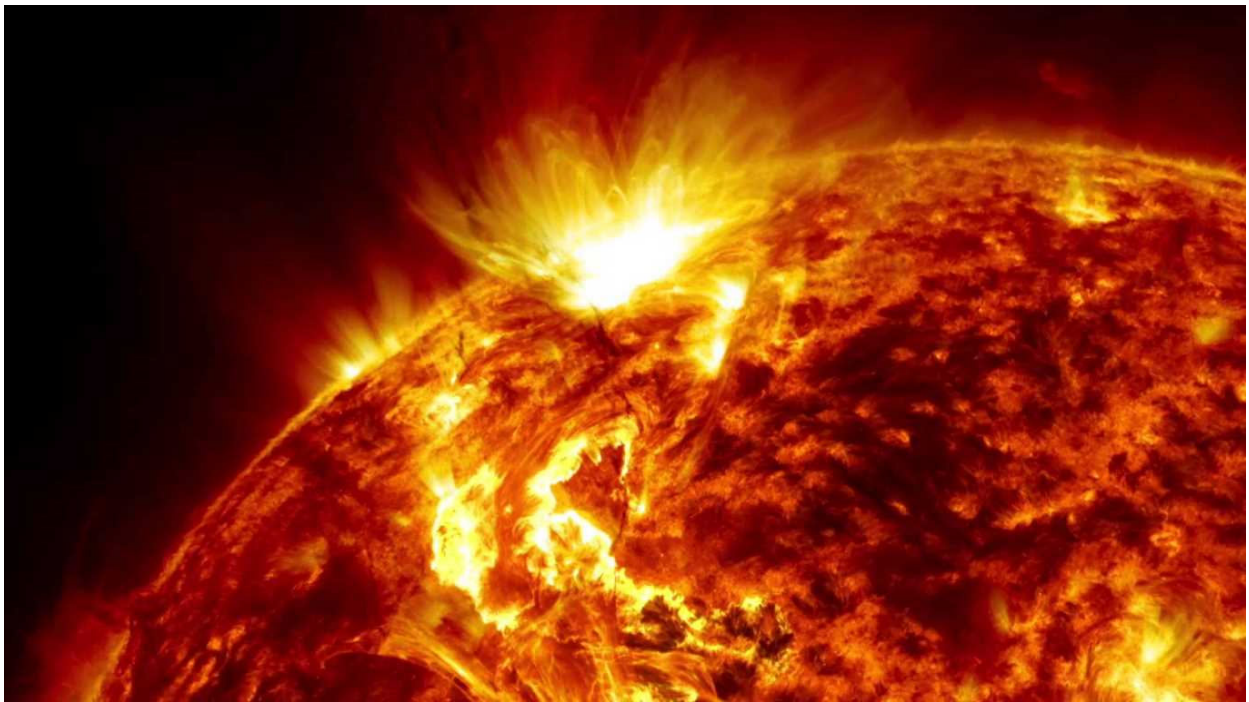
Action!

Solar Eclipse



Action!

Surface Features of the Sun



Action!

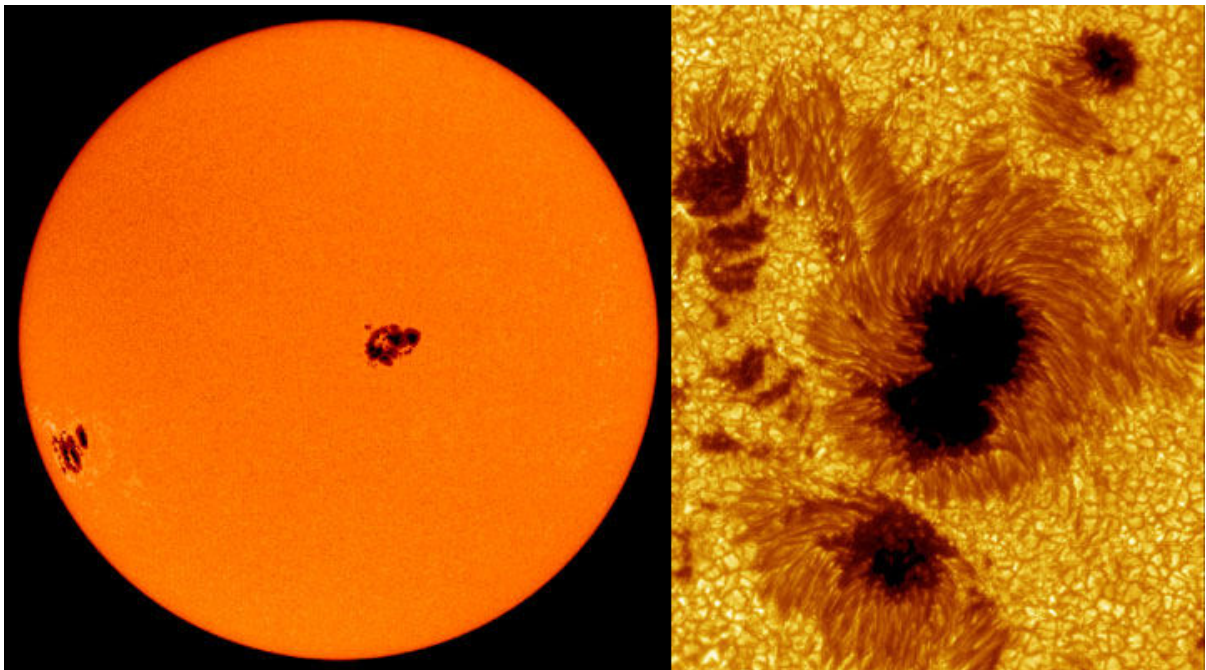
Sunspots

Appear as dark patches on the

surface and are cooler

than the surrounding areas.. They come and go and

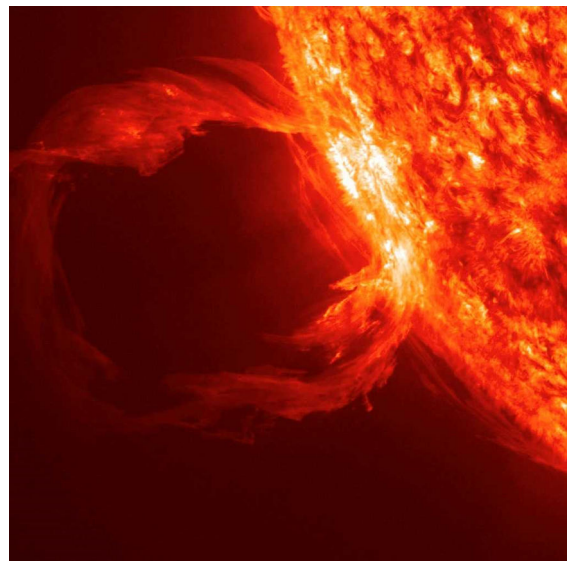
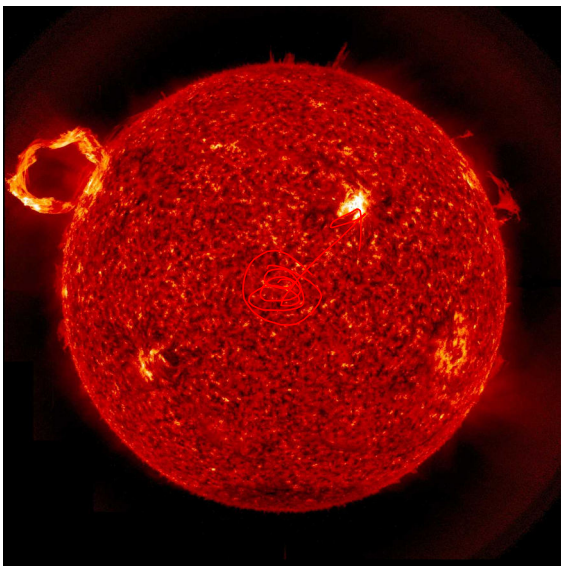
each is larger than Earth.



Action!

Prominences

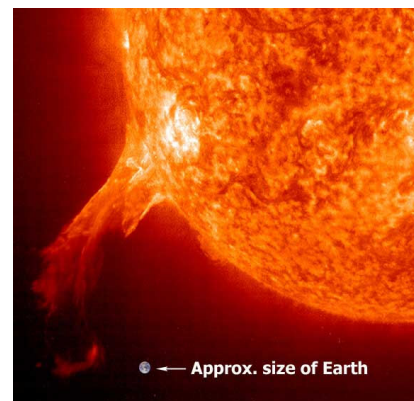
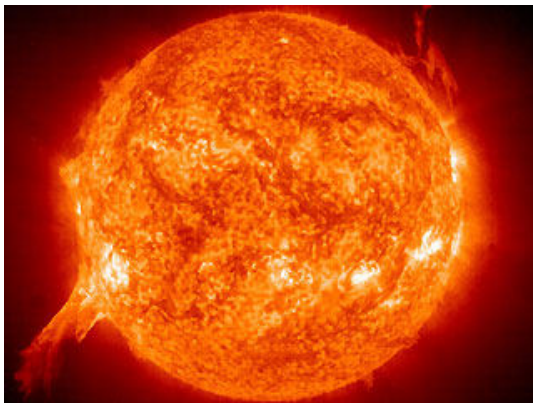
Large, curved stream of particles extending outward from the photosphere into the corona. They may last for many hours.



Action!

Solar Flares

Massive explosion of solar energy at Sun's surface. Solar flares fling hot plasma out into space and can damage orbiting satellites and electrical transmission lines on the ground.



Action!

Solar Flares

An extremely powerful solar flare is called a

coronal mass ejection.

These flares can reach Earth over the course of about

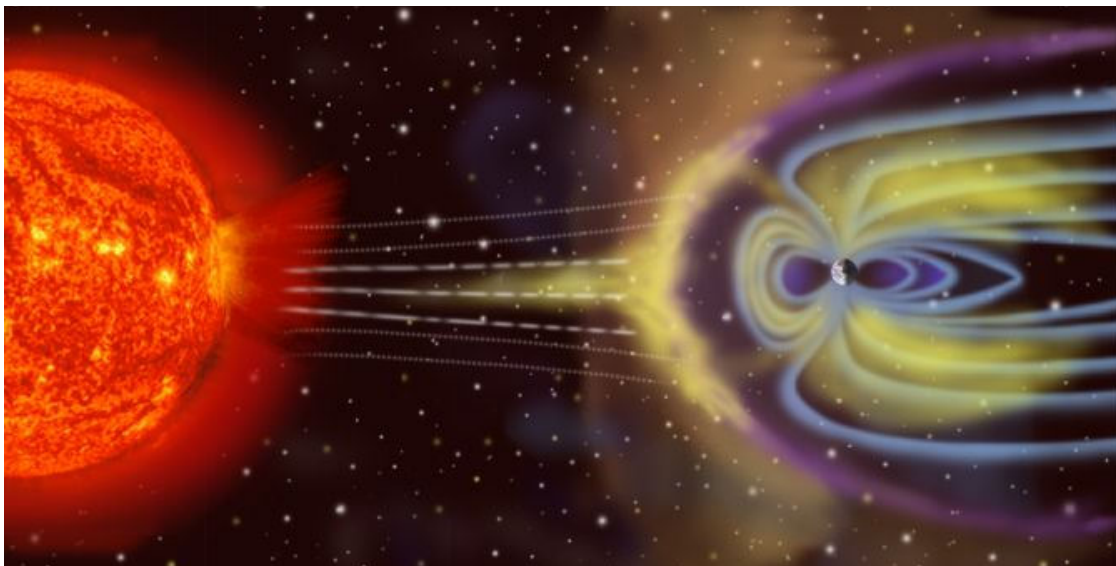
3 days where it meets Earth's

magnetic field. When the flare reaches the

field, the energy is diverted away from the planet's

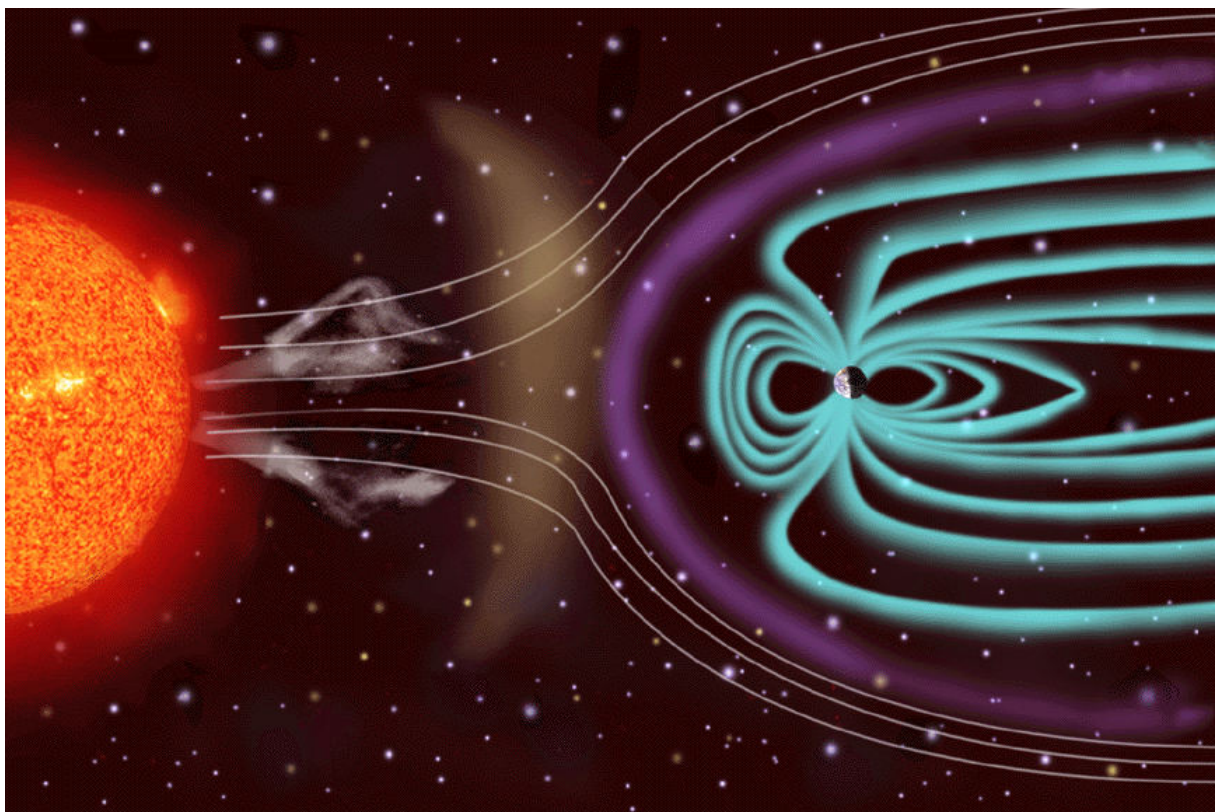
surface resulting in vivid and active

auroras.



Action!

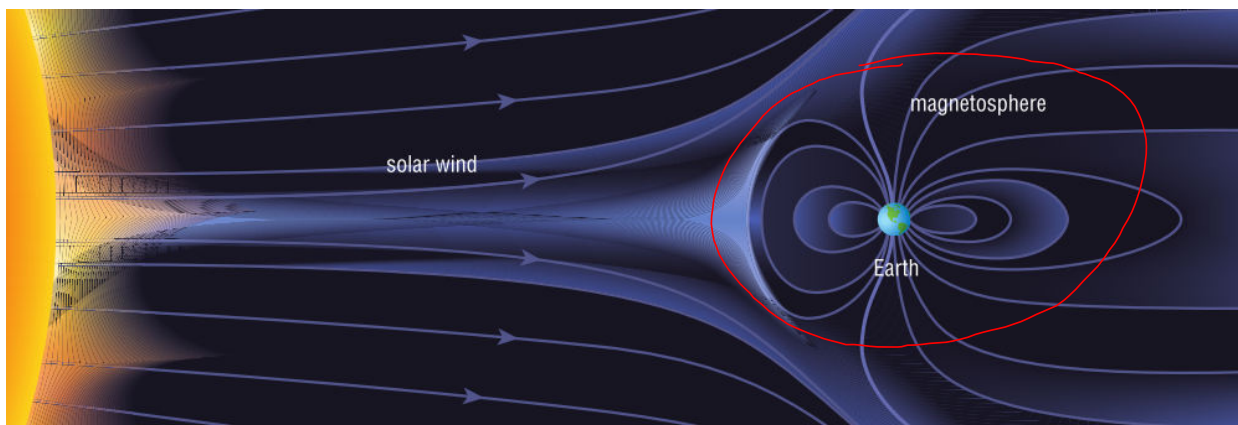
The Sun's Effects on the Earth



Action!

Solar Wind

The incredible amounts of heat at the surface of the Sun produces a steady stream of subatomic particles that emanates out from the Sun in all directions.



Action!

Aurora Borealis

Commonly known as the northern
lights, aurora borealis is created by the
solar wind, and results in displays of
green, yellow and red light.



Action!

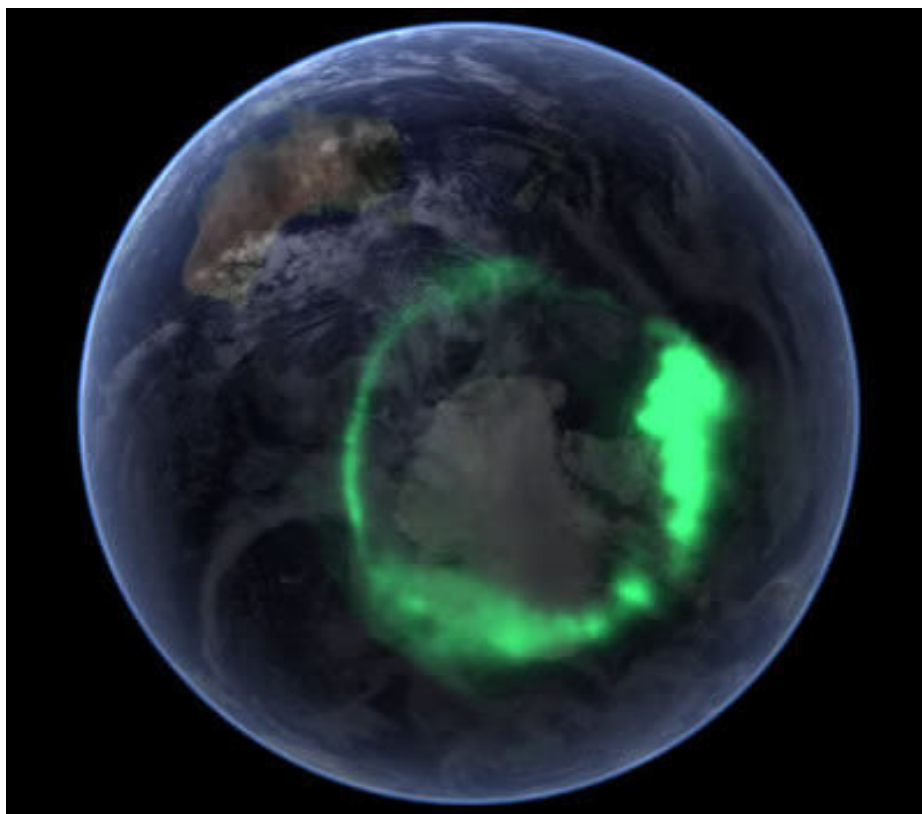
Aurora Borealis

The aurora borealis is produced when the charged particles of the solar wind collide with Earth's atmosphere. These particles are trapped by Earth's magnetic field and are swept toward the poles.



Action!

Aurora Borealis From Space!



Action!

Aurora Borealis From Space!



Action!

Aurora Borealis

From Space!



Consolidation

Fill It In, Fill It In

Group up with 1 or 2 other people and try and fill in the Sun note.

Homework!




















Fill in The Sun note

Pg. 321: 1 - 6

Add to your term table:

| core → radiative zone → convective zone
photosphere → chromosphere → corona
sunspot → prominence → solar flare
coronal mass ejection

Attachments

-  1 - Intro to Space - 1 - Contact Opening Scene.mp4
-  Intro to Space - The Beginning of the Universe.mp4
-  1 - Intro to Space - 1 - Celestial Objects.mp4
-  1 - Intro to Space - 2 - How Many.mp4
-  1 - Intro to Space - 4 - What Makes a Planet.mp4
-  1 - Z - Intro to Space - How Many Universes.mp4
-  A - Intro to Space - 1 - Contact Opening Scene.mp4
-  B1 - Stars - Star Size Comparison.mp4
-  B1 - (Stars) - Star Types.mp4
-  B1 - (Stars) - Massive Stars in the Milky Way.mp4
-  B1 - (Stars) - Tracking Stars Orbiting the Milky Way's Central Black Hole.mp4
-  B2 (The Solar System) - 5 Years of the Sun.mp4
-  B2 (The Sun) - 5 Years of the Sun.mp4
-  B2 (The Sun) - Solar Flare.mp4
-  B2 (The Sun) - Prominence.mp4
-  B2 (The Sun) - Corona.mp4
-  B2 (The Sun) - Solar Eclipse.mp4
-  B2 (The Sun) - Aurora Borealis.mp4
-  B2 (The Sun) - Aurora Borealis from Space.mp4